

# Stewart 8e 7.8.62: ideal gas speed distribution

$$\bar{v} = \frac{4}{\sqrt{\pi}} \left( \frac{M}{2RT} \right)^{3/2} \int_0^{\infty} v^3 e^{-\frac{Mv^2}{2RT}} dv$$

$M$  = molecular weight of gas  
 $R$  is the gas constant  
 $T$  is the gas temperature  
 $v$  is the molecular speed  
 $\bar{v}$  is the average speed of molecules in an ideal gas

$$\frac{Mv^2}{2RT} = \frac{v^2}{\left( \frac{2RT}{M} \right)} \equiv \frac{v^2}{V^2} \quad \text{where } V = \left( \frac{2RT}{M} \right)^{1/2}$$

must have dimensions of speed since the exponential must have an input that does not depend on the choice of units!

Therefore  $u = \frac{v}{V}$  is a dimensionless speed variable measuring speed in multiples of the velocity constant  $V$  which must characterize the velocity behavior of this system.

since there are no other parameters which can interfere, the average speed must be a "pure number" times  $V$ .

change of variable:  $u = \frac{v}{V} \quad du = \frac{dv}{V} \rightarrow v = Vu, \quad dv = V du, \quad v=0: u=0, \quad v \rightarrow \infty: u \rightarrow \infty$

both speed and differential of speed "scale" under this change of variable which acts like a unit conversion factor (inches to cm, ft-lbs to N-m, etc)

$$\bar{v} = \frac{4}{\sqrt{\pi}} V^{-3} \int_0^{\infty} (uV)^3 e^{-u^2} (V du) = \frac{4}{\sqrt{\pi}} V^{3+1-3} \left( \int_0^{\infty} u^3 e^{-u^2} du \right)$$

$$= \frac{4}{\sqrt{\pi}} V \left( \int_0^{\infty} u^3 e^{-u^2} du \right)$$

↑  
pure number  $\approx \frac{1}{2}$

comes speed units since result must be a speed.

$$\bar{v} = \frac{4}{\sqrt{\pi}} V \cdot \frac{1}{2} = \frac{2}{\sqrt{\pi}} V \approx 1.128 V$$

$$\text{or } \frac{\bar{v}}{V} = \frac{2}{\sqrt{\pi}} \approx 1.128$$

The average speed comes out equal to the "characteristic" speed  $V$  modulo a geometrical or numerical factor "of order unity"

Dimensional analysis is very important in applications.

Problem 7.8.66 on "stellar stereograph," is another good example where the distribution of mass in a star (or globular cluster) of radius  $R$  benefits from introducing a dimensionless <sup>radial</sup> variable  $u = r/R$  where  $r$  is the actual radius of a location in the star from its center. Instead of using a "dummy radial variable" in the integral, it introduces a copy variable  $s$ , so  $s = s/R$  is also appropriate, with  $0 \leq r \leq R \leftrightarrow 0 \leq u \leq R, 0 \leq s \leq R$ .

$$y(s) = \int_s^R \frac{2r}{\sqrt{R^2 - s^2}} x(r) dr \rightarrow y(r) = \int_r^R \frac{2t}{\sqrt{R^2 - t^2}} x(t) dt \quad \text{"dummy variable t"}$$

↑ same radial variable      ↑ can use its symbol here if we introduce